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Published in:
Optics Express

Link to article, DOI:
10.1364/OE.27.037877

Publication date:
2019

Document Version
Early version, also known as pre-print

Citation (APA):

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A Compact, Mobile, Low-Threshold Squeezed Light Source


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(Dated: September 4, 2019)

Strongly squeezed light finds many important applications within the fields of quantum metrology, quantum communication and quantum computation. However, due to the bulkiness and complexity of most squeezed light sources of today, they are still not a standard tool in quantum optics labs. We have taken the first steps in realizing a compact, high-performance 1550 nm squeezing source based on commercially available fiber components combined with a free-space double-resonant parametric down-conversion source. The whole setup, including single-pass second-harmonic generation in a waveguide, fits on a small breadboard and produces 9.3 dB of squeezing at a 5 MHz sideband-frequency. The setup is currently limited by phase noise, but further optimization and development should allow for a 19" sized turn-key squeezing source capable of delivering more than 10 dB of squeezing.

I. INTRODUCTION

Squeezed quantum states of light are a ubiquitous resource in numerous applications associated with quantum sensing, quantum communication and quantum computation [1,2]. One of the most celebrated examples is the application of squeezed light to improve the sensitivity of gravitational wave interferometers, thereby extending the volume in space within which gravitational events can be observed. [3–8]. A recent impressive improvement in observable volume is the eight-fold increase by the detection of 6 dB squeezed light in the gravitational wave detector GEO600 [9]. Quantum-enhanced sensitivity can also be achieved with squeezed light in tracking the motion and estimating bio-physical parameters of single living cells [10–12].

Apart from quantum sensing, squeezed light also has applications in quantum cryptography to extend the secure communication distance [13], to improve the cryptographic security [14] and to enable the implementation of quantum secure basic cryptographic primitives [15]. Finally, squeezed light has recently been found to be a viable resource for photonics continuous variable quantum computing due to development of new quantum error correcting codes [16,17] and due to the inherent scalability of the squeezed light source [4,15,21].

All of the above mentioned applications would naturally benefit from a compact, mobile and robust squeezed light source producing an appreciable amount of squeezing. However, in most experiments to date there have been a sharp trade-off between achieving high degrees of squeezing and the compactness (and thus robustness and transportability) of the source.

On one hand, quantum states have been significantly squeezed by up to 15 dB (that is, a reduction of 97% of the vacuum noise) using a nonlinear crystal embedded in an optical cavity [22,27], but the associated experimental setups have very large footprints and are not easily transportable due to the need for multiple mode-cleaning cavities for the pump and local oscillator to maximize the squeezing and cavity based second harmonic generation to supply the pump which easily exceeds 100 mW. Due to this immobility, the squeezed light sources are often built up around the application.

On the other hand, compact and mobile squeezed light sources have been constructed using either an optical waveguide in a single-pass configuration [28,29] or as a cavity [30], using a micron-sized Silicon Nitride ring cavity [31,32] or using a whispering gallery mode resonator [33], but in these systems the squeezing degree have been limited to maximum 3 dB. Using the optical Kerr effect in fibers, the production of 2.4 dB pulsed squeezed light was demonstrated on a mobile platform of 0.3 m² [34]. In all of these latter experiments, compactness has been traded with the squeezing degree.

The trade-off has been partially settled in a couple of realizations: The free-space squeezed light source – an optical parametric oscillator – used in the GEO600 gravitational wave detector has a footprint of 1.1 m × 1.3 m, a weight of approximately 70 kg, and produces about 10 dB squeezing [35]. Finally, a 50 cm × 60 cm free-space setup producing 6 dB of two-mode squeezed light has recently been demonstrated [36].

II. SETUP

In this article, we present the construction of a compact squeezed light source with a footprint of 30 cm × 45 cm producing 9.3 dB squeezing. While the source itself is a traditional double resonant parametric down-conversion source, we reduced its size by replacing bulky free-space optical components with less effective, but more compact commercially available fiber-alternatives and operate without the use of filter cavities. Only the couplings of the pump beam and the coherent control beam to the nonlinear cavity are obtained by free-space optics, as is the squeezed light output. In particular, the second-harmonic generator consists of a small single-pass waveguide module which pro-
Our experiment is presented in Fig. 1. The 1550 nm squeezed light is generated via parametric down-conversion by pumping a double resonant optical parametric oscillator (OPO) with light at 775 nm below threshold. The OPO itself, as sketched in the yellow inset, is a hemilithic cavity with a planar-convex periodically poled potassium titanyl phosphate (PPKTP) crystal as the non-linear medium. The crystal has a size of 9.8 mm × 2 mm × 1 mm, the planar facet is coated with an anti-reflection (AR) coating (R<0.1% @1550 nm, R<0.3% @775 nm), and the curved facet (radius of curvature (ROC) = 10 mm) is coated with a high-reflection (HR) coating (R=99.95% @1550 nm, R>99.5% @775 nm). The output mirror (ROC=25 mm) is coated with a partially reflective coating (R=90% @1550 nm, R=97.5% @775 nm). The cavity has a round-trip optical length of around 77 mm resulting in a free spectral range of approximately 3.9 GHz, a finesse of around 58 for the 1550 nm mode (full-width-half-maximum (FWHM) ≈ 66 MHz) and around 200 for the 775 nm mode (FWHM ≈ 17 MHz). The double-resonance is achieved by tuning the temperature of the crystal and the length of the cavity with a piezo-electric transducer.

An NKT Photonics E15 BOOSTIK 1550 nm fiber laser supplies the light to the setup. The light is divided into two paths: one serving as the local oscillator in the homodyne measurement setup. The other path is further split into a pumping path and a path for the coherent control beam [37], in the following called the pilot beam. The pumping path starts with a 1550 nm electro-optic modulator (EOM) to modulate the phase at 100 MHz for the OPO Pound-Drever-Hall (PDH) lock. The EOM has a maximum input power of 300 mW and an insertion loss of 3 dB. This limits the input power into the LiNbO3 second-harmonic generator (SHG) wave-guide module
FIG. 2. Overlapped Allan deviation (OADEV) and power-spectral density of polarization noise measurements in the local oscillator path. The data was taken at a 1 Hz sampling rate and is normalized to the mean and has mean subtracted. The blue trace is without any polarizing components. The yellow trace is with one polarization sensitive isolator and the green trace is with two isolators. The red trace is polarization insensitive amplitude noise.

(NTT Electronics WH-0775-000-F-B-C) to 150 mW, resulting in around 8 mW of 775 nm light to be used to pump the OPO.

The pilot path implements a coherent-locking scheme for locking the relative phase between the pump light and the local oscillator by using a fiber acousto-optic modulator (AOM) to up-shift the frequency of a 1550 nm beam by 40 MHz. This pilot field enters the OPO through the HR side and interacts with the pump field through difference-frequency generation. The reflected light is detected and down-mixed with an 80 MHz tone to generate an error signal for locking the phase of the pilot field to the pump field. The locked pilot field is transmitted with the squeezing and beats with the local oscillator. After detection in the homodyne detector, it is down-mixed with a 40 MHz tone to provide an error-signal for locking the phase between the local oscillator and the squeezed field.

Both the pump and the pilot beams are coupled out of the fibers and collimated by fiber collimators. Using two mode-matching lenses and two steering mirrors, each of the free-space beams are coupled into the OPO. All free-space optics are placed on the 30 cm × 45 cm breadboard which leaves enough space to place all fiber components on it as well. In the actual experiment, the fiber components were not attached to the breadboard out of convenience, but will be placed in the box in the final version of the device.

The squeezed light is characterized in a balanced homodyne setup which is placed on a neighboring breadboard and uses a bright (10 mW) 1550 nm beam as a local oscillator. The overlap between the squeezed light and the local oscillator is optimized by coupling both fields into a triangular reference cavity. This allows us to achieve a fringe visibility of around 99%. The homodyne detector uses InGaAs photodiodes (ηQE > 99%). The photo-electric signal is analyzed with a spectrum analyzer.

III. RESULTS

The setup utilizes polarization maintaining (PM) fiber components, and every time two PM fibers are combined in a mating sleeve, a small polarization mismatch can appear due to an imperfect slow-axes alignment. With many components, such a mismatch can build up and result in quite severe polarization noise. This problem can be circumvented by using a polarization stabilizing feedback loop, but a simpler solution is to insert polarization filtering components along the path, thus preventing polarization mismatch to build up. The downside to this simple solution is the introduction of some excess amplitude noise.

In our setup, the EOM, AOM, SHG and circulators are all polarization filtering components thus preventing mismatch to build up in the pump and pilot paths. For the local oscillator, however, polarization noise was a problem, and fiber isolators were therefore inserted into the fiber path.

We characterize the effect of inserting polarizing components by inserting a power-meter after a polarizing beam-splitter (PBS) measuring at a sampling rate of 1 Hz. Fig. 2 shows a plot of the overlapped Allan deviation and the power spectral density of these measurements. We measured without any isolators, with isolator 1 and with two isolators, isolators 1 and 2. For their positions see the experimental schematic in Fig. 1. As a baseline, we measure the polarization insensitive amplitude noise. The data is normalized to the mean, and
the mean is subtracted in order to get the fluctuations relative to the mean.

From the Allan deviation, we see a large increase in long-term stability of the power when isolator 1 is present. This can also be seen in the power spectral density, with a difference of almost two orders of magnitude between having no polarizing components and having isolator 1. Between having one and two isolators, the difference is negligible for both graphs which is to be expected as no additional fiber components are present after isolator 2. By comparing the green/orange traces with the red trace, it seems that the amplitude noise with isolators is in general a bit higher than the intrinsic noise from the laser. This could be due to small polarization mismatches being converted to amplitude noise by the isolators.

The performance of our setup is first characterized by estimating the pump threshold power via a classical gain measurement of a 1550 nm field interacting with the pump field in the crystal. For this measurement, the fiber AOM is removed and the transmitted power of the 1550 nm field is measured while varying the input power of the 775 nm pump field. The gain is then estimated by comparing with the transmitted 1550 nm light when the pump is blocked. The calculated gain values are plotted in Fig. 3. The gain is modelled as

$$g = \frac{1}{\left(1 - \sqrt{\frac{P_{p}}{P_{thr}}}\right)^2},$$

where $P_{p}$ is the input pump power and $P_{thr}$ is the OPO threshold power. From the fit of the experimental gain values, we can extract an OPO threshold power of $5.12 \pm 0.03$ mW. This shows that even though the system only has 8 mW of pump power available due to lossy fiber components, the double resonance allows the squeezer to utilize its full potential, since squeezing is best generated below threshold.

In order to characterize the squeezing performance, we first look at the 5 MHz side-band frequency using a resolution-bandwidth of 300 kHz and video-bandwidth of 300 Hz. We measure for 200 ms and average each trace 100 times. The pump power is varied in the range 0–4 mW, and the local oscillator phase is locked to squeezing and anti-squeezing, respectively. The measured values as a function of pump power are shown in Fig. 4, corrected for electronic noise, which is about 22 dB below shot noise.

The squeezing increases with power until a maximum of around 9.3 dB below shot noise at around 2.5 mW pump power. After this, the squeezing begins to degrade with increasing pump power, suggesting that the measurement suffers from a significant amount of phase noise. The anti-squeezing is almost insensitive to phase noise and increases for all pump powers. The full power dependence of the (anti-)squeezing variance ($\langle \delta X_+^2 \rangle$) including phase noise can be modelled as (normalized to shot noise)
FIG. 4. a) Squeezing and anti-squeezing relative to shot noise as a function of pump power at a side-band frequency of 5 MHz. The blue points are squeezing and the orange points are anti-squeezing. The theoretical model of Eq. 2 is fitted to the data and plotted in solid lines. The purple dashed line is the fitted squeezing model in the absence of phase noise. The electronic noise, which is 22 dB below shot noise, was subtracted from the data. The error-bars assume a 5% error on the power. b) Spectrum of the squeezing and anti-squeezing from 1–120 MHz for different pump powers. The traces are corrected for electronic noise and normalized to the shot noise. The thin solid lines are fits of Eq. 2. Bands around 40 MHz, 80 MHz and 100 MHz are excluded from the fit as they contain the 40 MHz up-shifted pilot tone and electronic pick-up of modulation signals, and the 80 MHz and 100 MHz peaks are removed from the figure.

the way out beyond 100 MHz. Once again, for high pump powers the squeezing starts to degrade at low frequencies due to phase noise, and from the fits of the curves, we extract similar values for the efficiency and phase noise. A curious exception is the 3.5 mW trace that surprisingly reaches all the way to 9.6 dB below shot noise for frequencies of 1–7 MHz and only contains about 12 mrad of phase noise from the fit.

IV. SUMMARY

In summary, we have demonstrated that a squeezed light source based on a traditional bulk squeezing cavity can be built with a small footprint on an area smaller than what is available in standard 19 inch boxes. We achieved this by replacing most free-space optics with fiber optics except for the in- and out-coupling of the OPO. The result was in particular made possible by a single-pass waveguide SHG module which provided enough pump power for the low threshold of only a couple of milliwatts, which in turn was due to the double-resonance of the OPO.

Further engineering on setup packaging will enable even smaller foot prints. A mobile turn-key squeezed laser in a transportable box might make squeezing a standard tool in quantum optics labs, which are ready to explore many yet undiscovered applications.

ACKNOWLEDGEMENTS

The authors acknowledge support from the Innovation Fund Denmark through the Quantum Innovation Center, Qubiz as well as support from the Danish National Research Foundation, Center for Macroscopic Quantum States (bigQ, DNRF142). Furthermore the authors acknowledge the EU project CiViQ (grant agreement no. 820466) and the EU project UNIQORN (grant agreement no. 820474).

squeezed states of light. [Nature Photonics 7, 613 (2013)]


